

# Radio luminosity functions

## with machine learning and Radio Galaxy Zoo

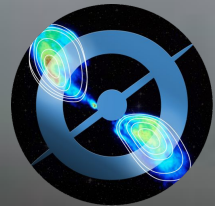
Matthew Alger (ANU/Data61)

O. Ivy Wong (ICRAR/UWA)

Cheng Soon Ong (Data61/ANU)

Naomi McClure-Griffiths (ANU)

Slides: <http://www.mso.anu.edu.au/~alger/asa-19>



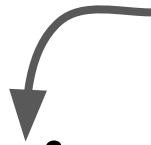
Australian  
National  
University



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## with machine learning and Radio Galaxy Zoo

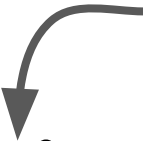
Comoving density of radio sources  
as a function of radio luminosity



# Radio luminosity functions

with machine learning and Radio Galaxy Zoo

Comoving density of radio sources  
as a function of radio luminosity



# Radio luminosity functions

## with machine learning and Radio Galaxy Zoo




Approximating functions  
based on existing data

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Comoving density of radio sources  
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Approximating functions  
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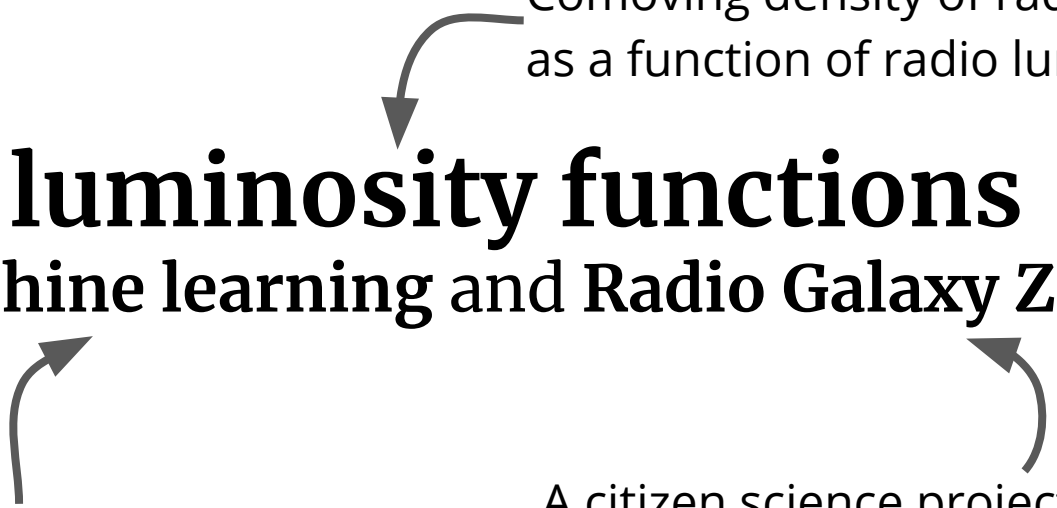


A citizen science project for  
matching radio emission to  
infrared galaxies



# Radio luminosity functions with machine learning and Radio Galaxy Zoo

Comoving density of radio sources  
as a function of radio luminosity

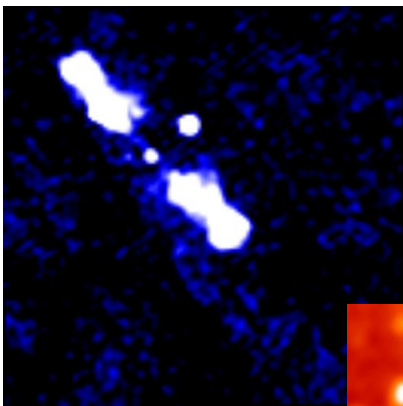
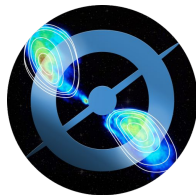


Approximating functions  
based on existing data

A citizen science project for  
matching radio emission to  
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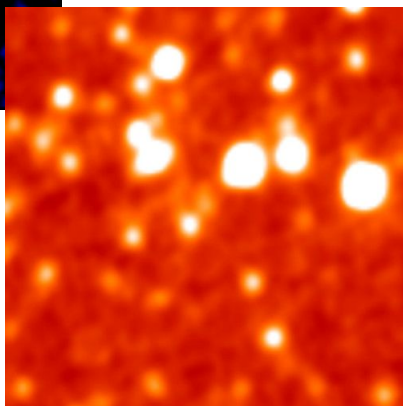
*Use machine learning to approximate how citizen scientists cross-identify, then cross-identify everything and make a luminosity function with a huge sample size*

# Radio Galaxy Zoo

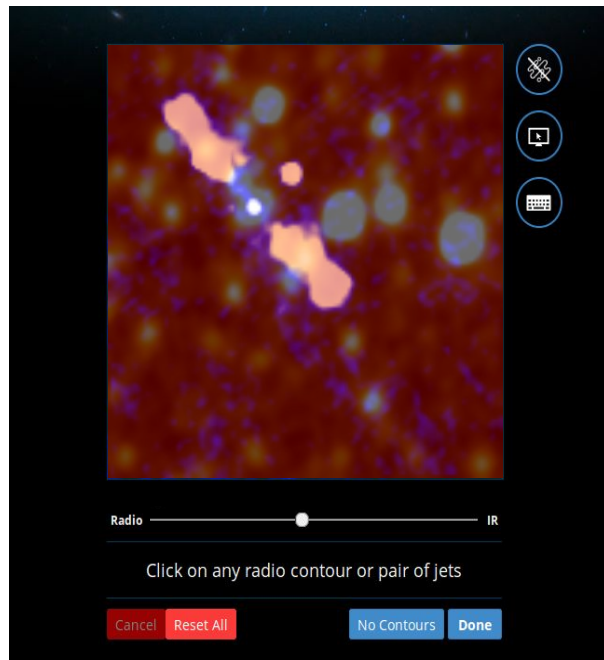


1.4 GHz radio  
(FIRST)

3.4  $\mu\text{m}$  infrared  
(WISE)

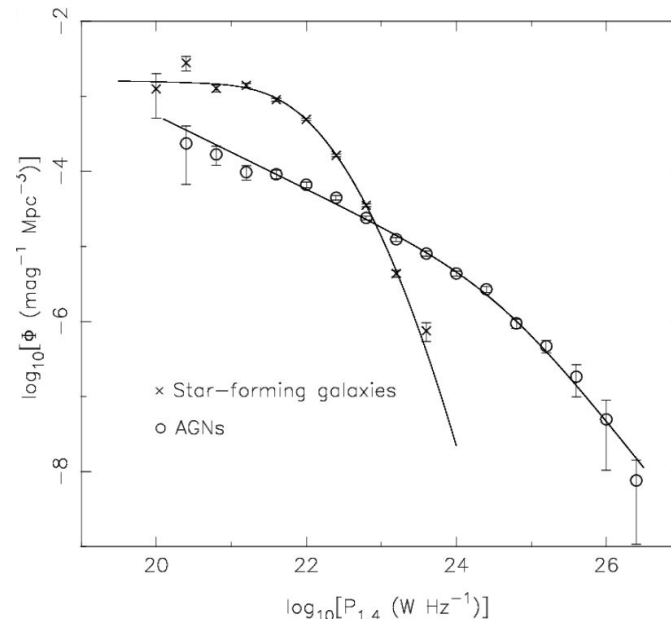


Citizen scientists cross-identify radio emission with infrared host galaxies



# Radio luminosity functions

- Comoving density of radio sources as a function of radio luminosity
  - Units of  $\text{dex}^{-1} \text{ Mpc}^{-3}$
  - Distribution of radio source luminosities in a *physically meaningful* way
- Fractional radio luminosity functions
  - Luminosity distribution of physically-selected subsets may be different
  - Helps understand evolution and structure of radio galaxies



Radio luminosity function divided into radio due to star formation and radio due to active galactic nuclei.

*Image: Mauch & Sadler (2007)*



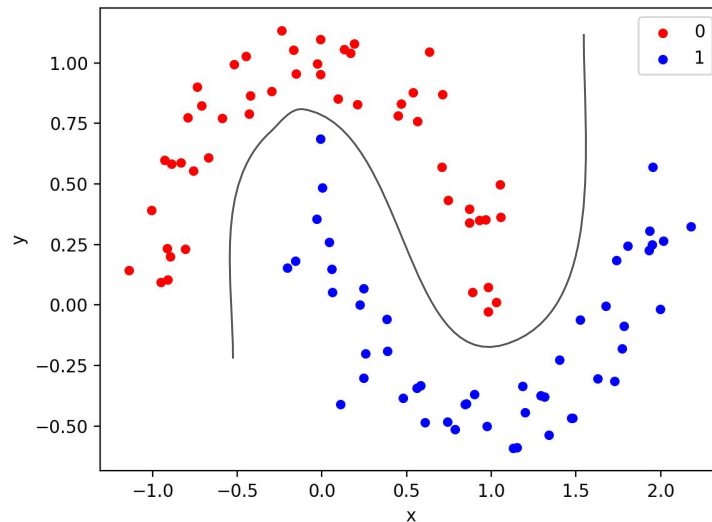
# Binary classification

- Find a function that separates objects into two classes
- Well-understood

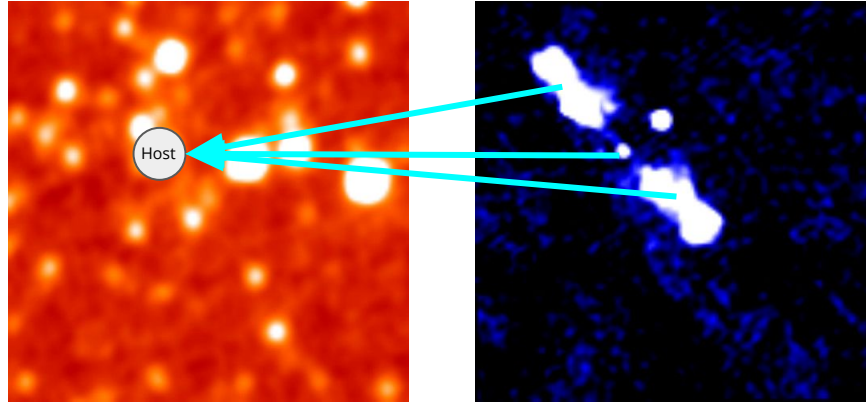
Equivalent:

$$\left\{ \begin{array}{l} f: \mathbb{R}^d \rightarrow \mathbb{R} \\ g: \mathbb{R}^d \rightarrow [0, 1] \\ h: \mathbb{R}^d \rightarrow \{\top, \perp\} \end{array} \right.$$

$h(x) = g(x) > 0$   
 $g(x) = \sigma(f(x))$

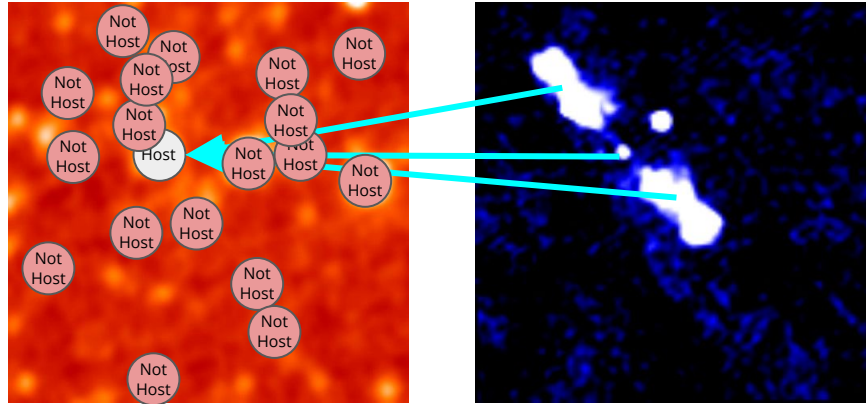


# Learning from Radio Galaxy Zoo



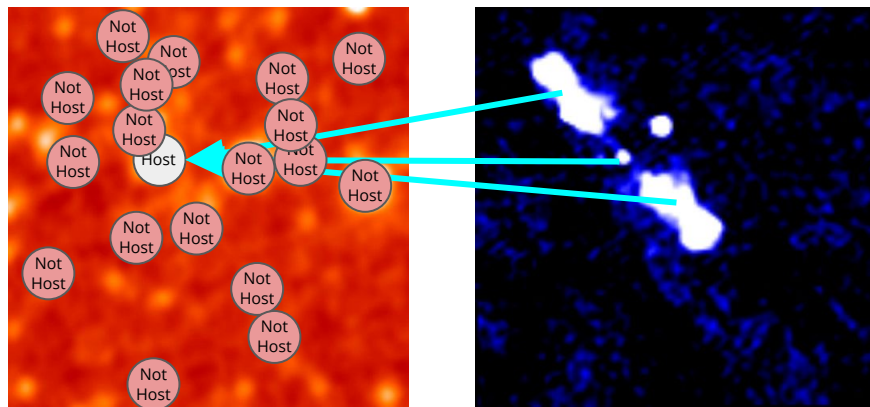
- Assign hosts positive labels

# Learning from Radio Galaxy Zoo

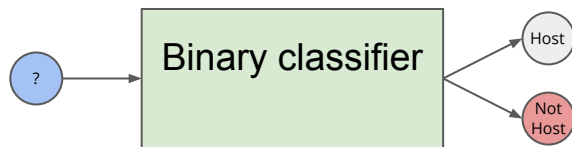


- Assign hosts positive labels
- Assign everything else negative labels

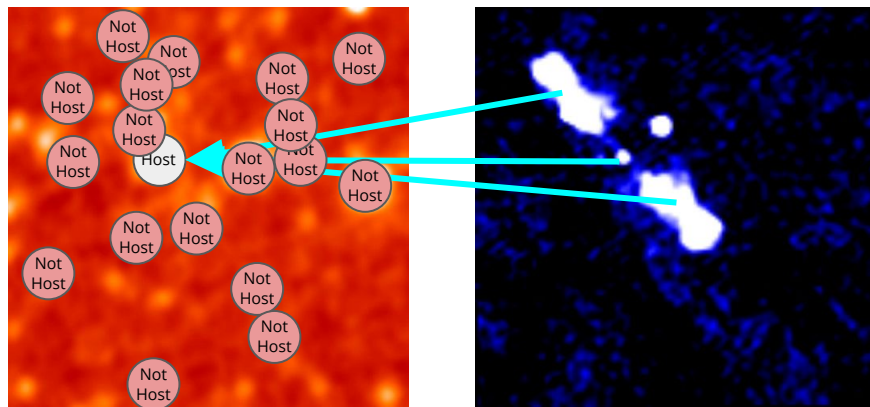
# Learning from Radio Galaxy Zoo



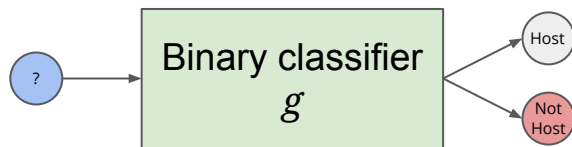
- Assign hosts positive labels
- Assign everything else negative labels
- Train classifier to identify *host* and *not host* classes



# Learning from Radio Galaxy Zoo



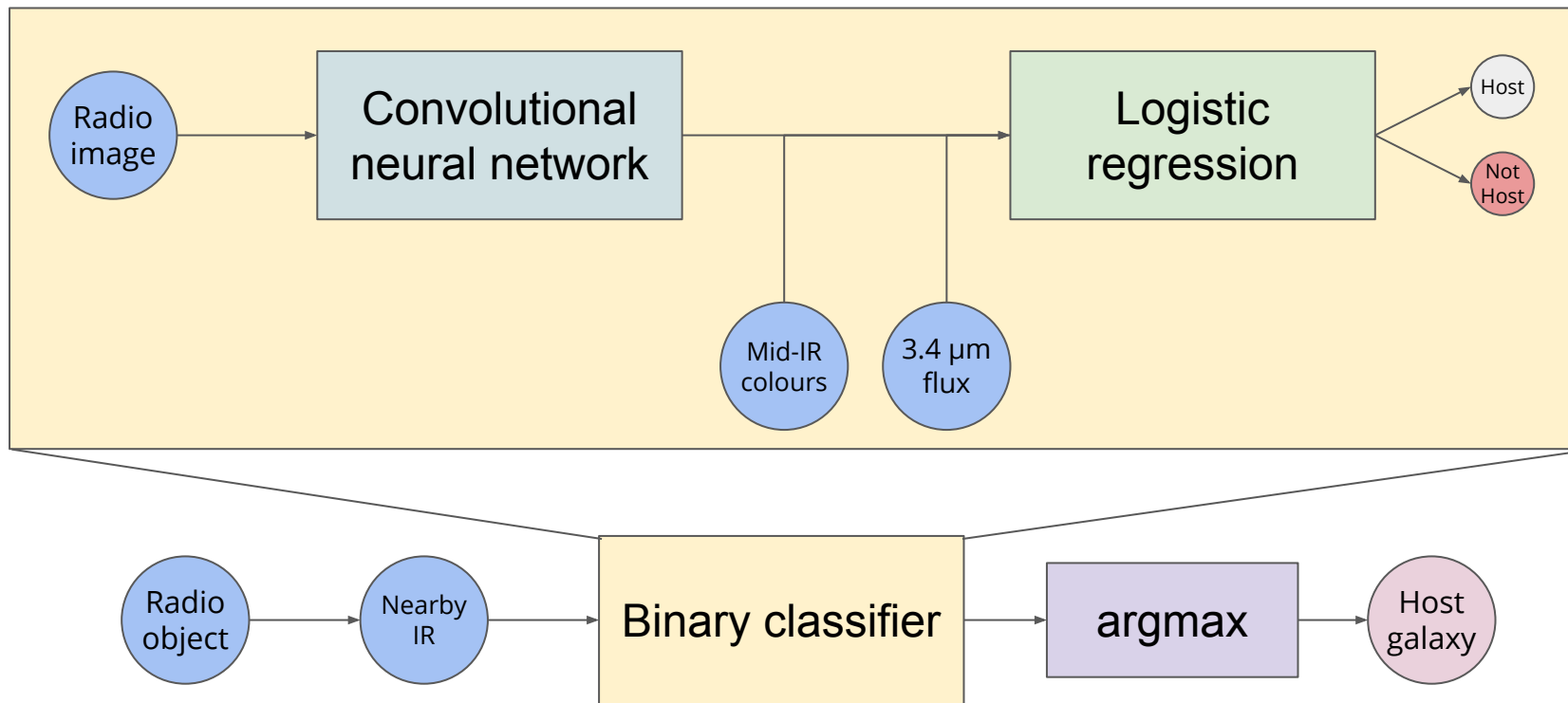
- Assign hosts positive labels
- Assign everything else negative labels
- Train classifier to identify *host* and *not host* classes



$\text{xid} : \text{Radio} \rightarrow \text{IR}$

$$\text{xid}(\mathbf{r}) = \underset{i \in \text{IR objects}}{\operatorname{argmax}} g(i) \mathcal{N}(\mathbf{r}, i)$$

# Binary classification model



# Luminosity function

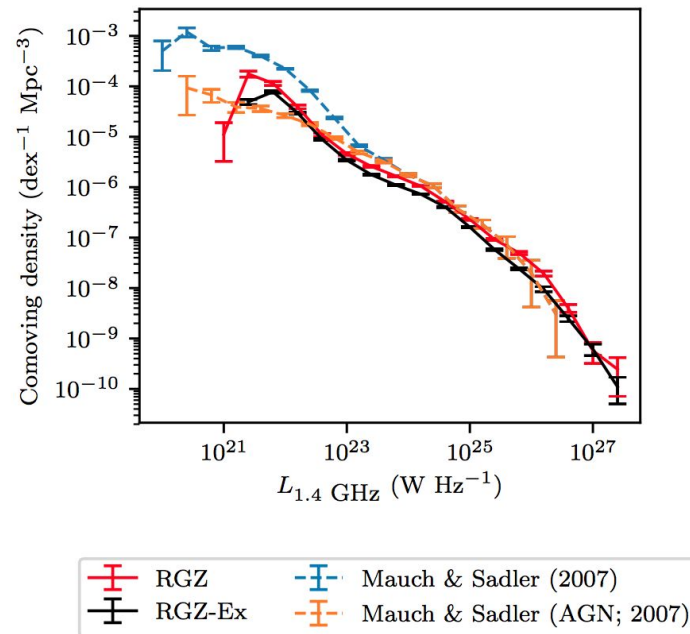
- RGZ-Ex contains 157 007 cross-identified radio sources with 30 743 redshifts
- Large sample allows us to build a radio luminosity function of extended sources
  - Luminosities up to  $10^{27}$  W/Hz
  - Close match to Mauch and Sadler (2007) radio AGN luminosity function

## Radio Galaxy Zoo: radio luminosity functions of extended sources

M. J. Alger<sup>1,2\*</sup> et al.

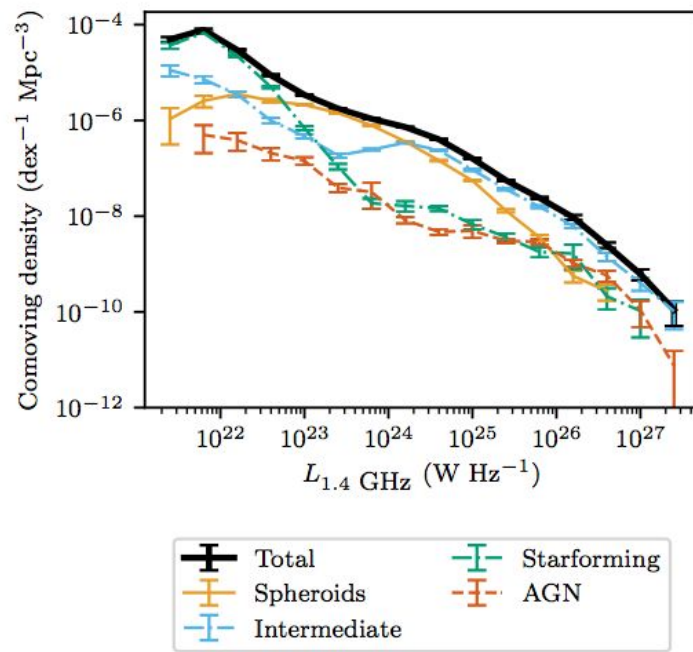
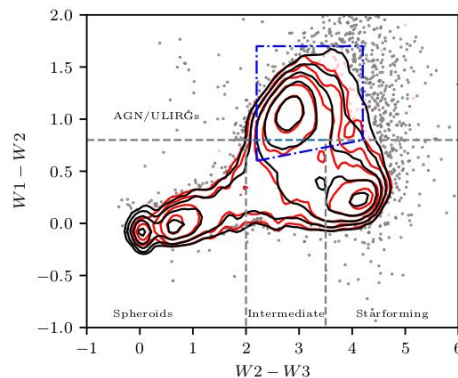
<sup>1</sup>Research School of Astronomy and Astrophysics, The Australian National University, Canberra, ACT 2611, Australia

<sup>2</sup>Data61, CSIRO, Canberra, ACT 2601, Australia



# Fractional luminosity function (Mid-IR)

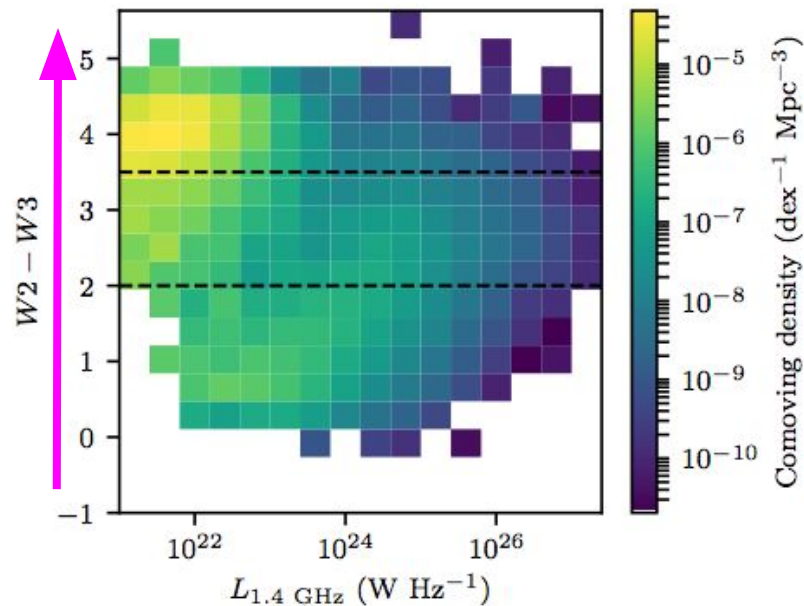
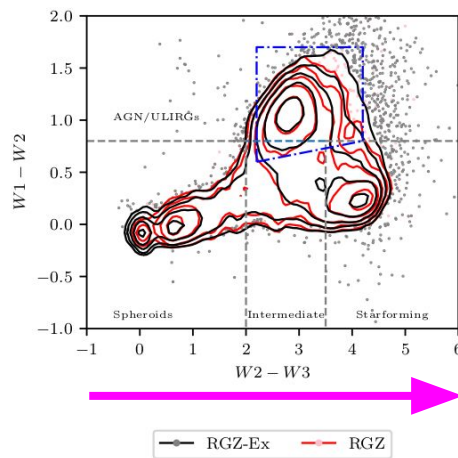
- Divide radio luminosity function based on mid-infrared host colours
  - “Extended” star-forming sources below  $10^{23}$  W/Hz (visually verified)
  - Radio-loud sources dominated by “intermediate” galaxies





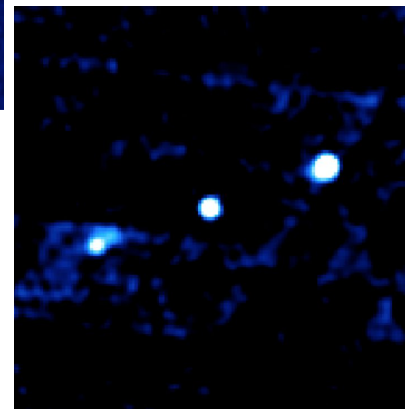
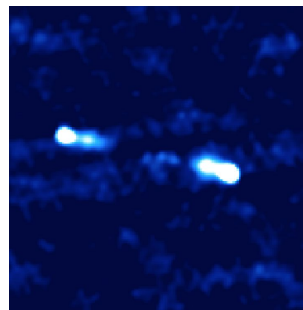
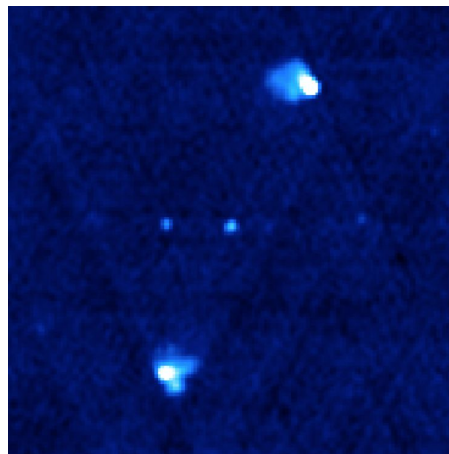
# Bivariate luminosity function (Mid-IR)

- Divide radio luminosity function based on 12  $\mu\text{m}$ /4.6  $\mu\text{m}$  colour
  - “Extended” star-forming sources below  $10^{23}$  W/Hz (visually verified)
  - Radio-loud sources dominated by “intermediate” galaxies



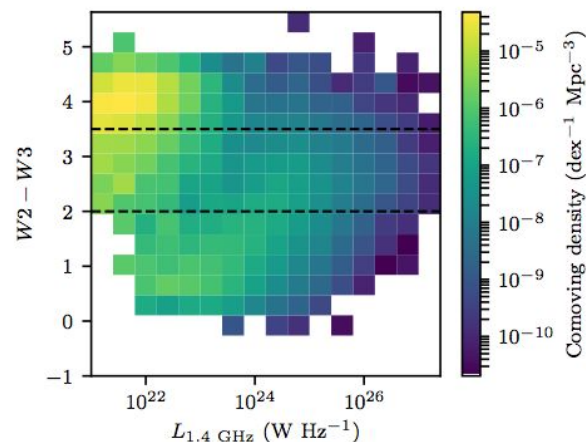
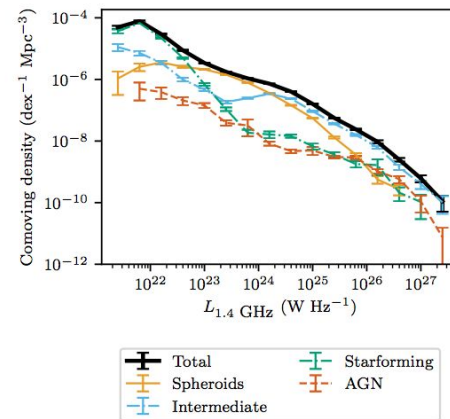
# The RGZ-Ex catalogue

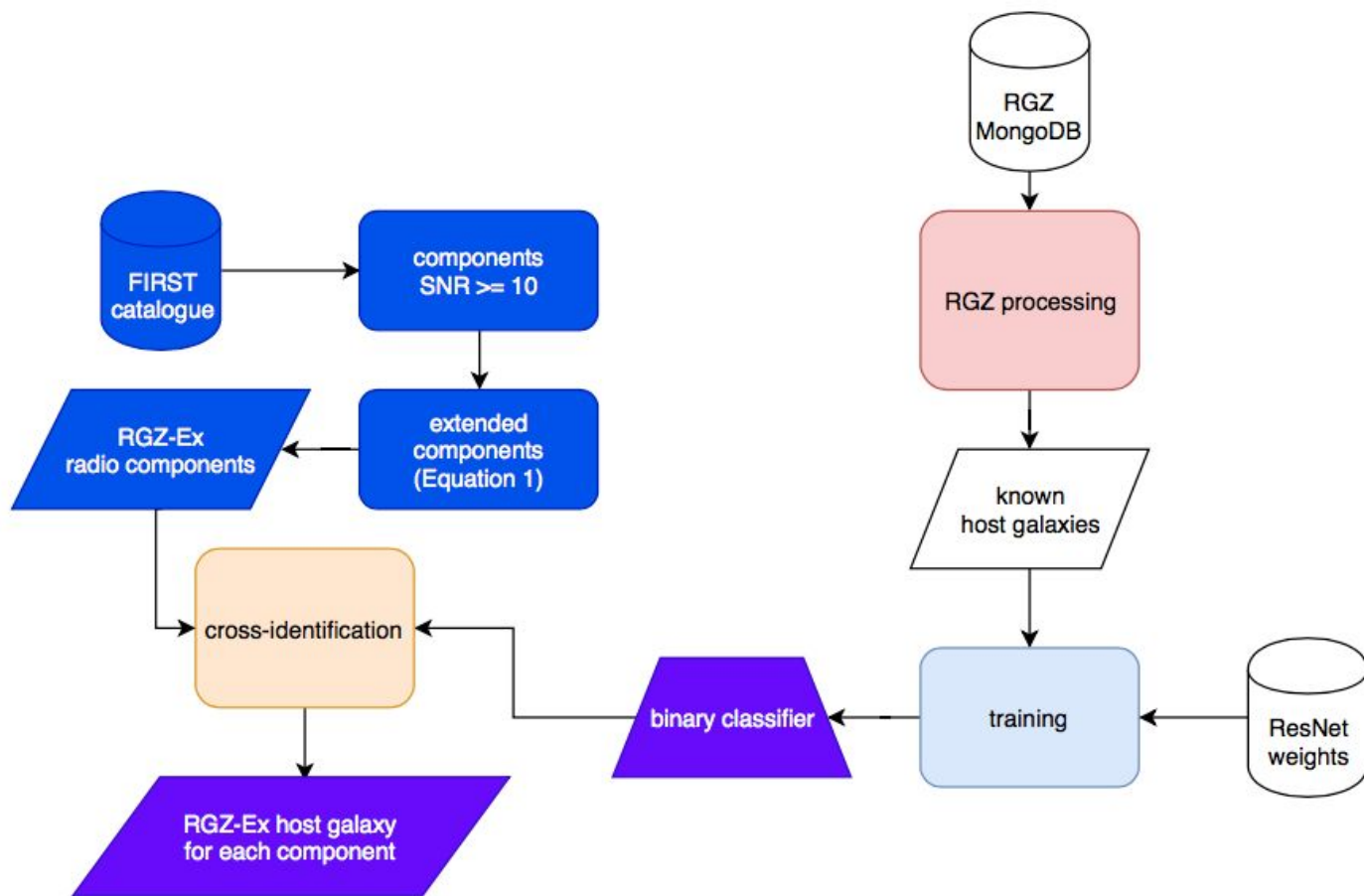
- Catalogue of 157 007 candidate radio sources and their hosts
- Large but noisy
- Contains around *sixty* previously unidentified giant radio galaxies ( $\geq 1$  Mpc)

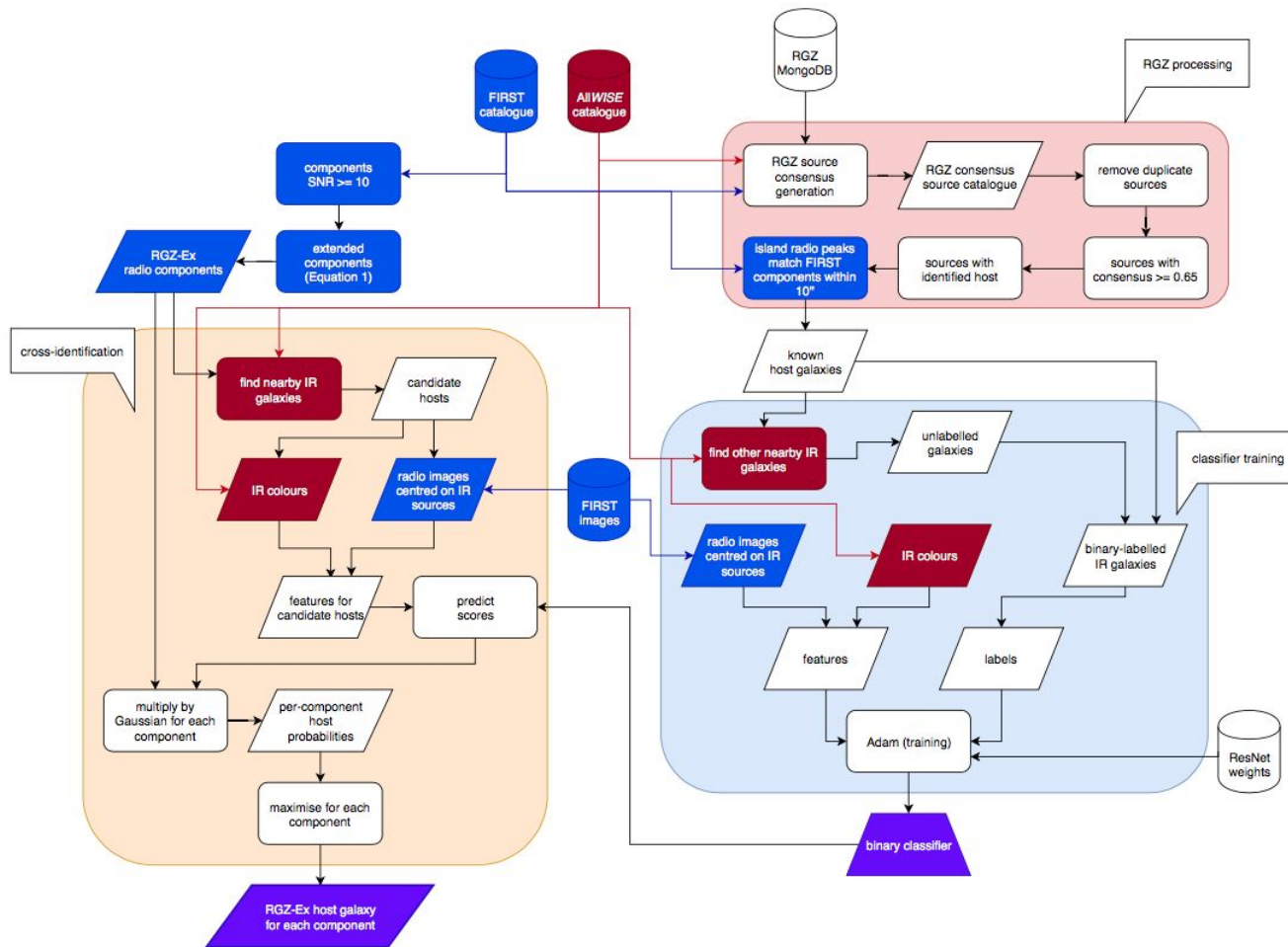


# RGZ-Ex and luminosity functions

- We developed an automated, machine learning approach to radio-infrared cross-identification
- We created a huge catalogue of candidate radio sources and their hosts
- We estimated fractional radio luminosity functions of extended radio sources
- We found ~60 new giants

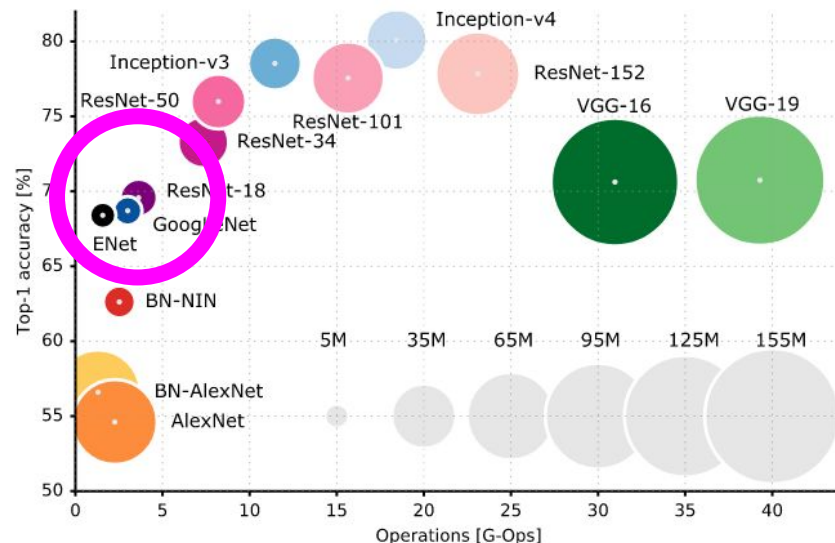






# Binary classification model

- ResNet-18 (multiclass)
  - Good accuracy
  - Low complexity
  - Very fast to train and use
- Remove last layer and replace with a binary classifier
- Add non-image features
  - Mid-infrared colours
  - 3.4  $\mu\text{m}$  flux
  - Room for improvement — e.g. add redshifts



Trade-offs between network complexity and accuracy on ImageNet.

Image: Canziani et al. (2016)

# Cross-identification as binary classification

$$\begin{aligned} \text{xid} : \text{Radio} &\rightarrow \text{IR} \\ \text{xid}(r) &= \underset{i \in \text{candidate IR hosts}}{\operatorname{argmax}} f(i; r) \end{aligned}$$

$$\begin{aligned} &\text{where} \\ &f : \mathbb{R}^d \rightarrow \mathbb{R} \\ &f(i) = p(\text{host} \mid i) \\ &\text{is a binary classifier} \end{aligned}$$