## Radio luminosity functions with Radio Galaxy Zoo and machine learning

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Slides: http://www.mso.anu.edu.au/~alger/sparcs-ix



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## Radio luminosity functions

- Comoving density of radio sources as a function of radio luminosity
  - Units of dex<sup>-1</sup> Mpc<sup>-3</sup> or mag<sup>-1</sup> Mpc<sup>-3</sup>
  - Comoving density accounts for universe size and shape over cosmic time
  - Distribution of radio source luminosities in a *physically meaningful* way
- Fractional radio luminosity functions
  - Luminosity function of a subset of sources
  - Luminosity distribution of physically-selected subsets may be different
  - Helps understand evolution and structure of radio galaxies



Radio luminosity function divided into radio due to star formation and radio due to active galactic nuclei. *Image: Mauch & Sadler (2007)* 

### Radio Galaxy Zoo

FIRST Survey Northern Sky Coverage, 2014 December 17



FIRST (radio)



WISE (infrared)



Radio Galaxy Zoo



Zooniverse ID (3)	RA (4)	Declination (5)	N <sub>votes</sub> (6)	$\frac{N_{\text{total}}}{(7)}$	CL (8)
ARG000255v	251.679244	23.382107	41	42	0.98
ARG000255x	163.799660	23.384972	58	58	1.00
ARG000255y	138,960429	23.381641	43	43	1.00
ARG000255z	126.215156	23.381729	35	35	1.00
ARG0002560	149.273620	23.381661	40	40	1.00
A P C0002561	167 047509	22 291620	24	25	0.07

Radio Galaxy Zoo cross-identification catalogue

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## Cross-identification as binary classification

#### • Cross-identification

- Match radio emission to infrared host galaxies
- Output of Radio Galaxy Zoo
- Can be cast as binary classification
  - Binary classification is well-understood
  - Lots of off-the-shelf classification models
  - Easy to train
- Problems:
  - Converting cross-identification catalogues to binary labels loses information
  - Unclear how uncertainties in this formulation are related to dataset or physical uncertainties



 $\begin{array}{l} xid: Radio \rightarrow IR\\ xid(r) = \underset{i \in IR \ objects}{argmax} \ f(i; r) \end{array}$ 



 Assign hosts positive labels

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- Assign hosts positive labels
- Assign everything else negative labels



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- Assign hosts positive labels
- Assign everything else negative labels
- Train classifier to identify host and not host classes





- Assign hosts positive labels
- Assign everything else negative labels
- Train classifier to identify *host* and *not host* classes



 $\begin{array}{l} \text{xid}: \text{Radio} \rightarrow \text{IR} \\ \text{xid}(r) = \underset{i \in \text{IR objects}}{\operatorname{argmax}} g(i) \ \mathcal{N}(r, i) \end{array}$ 



## Binary classification model

- ResNet-18 (multiclass)
  - Good accuracy
  - Low complexity
  - Very fast to train and use
- Remove last layer and replace with a binary classifier
- Add non-image features
  - Mid-infrared colours
  - 3.4 µm flux
  - Room for improvement e.g. add redshifts
- Accurate for ~96% of sources



Trade-offs between network complexity and accuracy on ImageNet. Image: Canziani et al. (2016)



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## Binary classification model



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### Luminosity function of extended sources

- RGZ-Ex contains 214 214 cross-identified radio components with 26 268 redshifts
  - >4x more components than RGZ
  - >2x more components with redshifts than RGZ
- Large sample allows us to build a radio luminosity function of extended sources
  - $\circ$  Luminosities 10<sup>21</sup>–10<sup>27</sup> W/Hz



## Fractional luminosity function (Mid-IR)

- Divide radio luminosity function based on mid-infrared host colours
  - "Extended" star-forming sources below 10<sup>23</sup> W/Hz
  - Radio-loud sources

dominated by "intermediate" galaxies





#### Fractional luminosity function (Mid-IR)



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## Source finding with cross-identification

- We matched 214 214 radio components to 157 007 host galaxies
- If two components have the same host they are part of the same source
  - Many false positives
- RGZ-Ex candidate source catalogue
  - 157 007 candidate extended sources
  - All data required to reproduce our results
  - At least 10 new giants



Two giant radio galaxies identified in our catalogue, each 1.1 Mpc across. Images: FIRST



# Looking forward: EMU

- Very different results to previous work on EMU-ATLAS
  - Much lower source density on sky
  - Higher angular resolution more extended sources
  - More training and prediction data
  - Alger+18, <u>doi:10.1093/mnras/sty1308</u>
- Generalisation to EMU will be non-trivial
- Radio Galaxy Zoo for EMU?



ATLAS observations of CDFS. Image: ATLAS DR3, Franzen+2015

